## **COMMENTS**

Comments are short papers which criticize or correct papers of other authors previously published in the Physical Review. Each Comment should state clearly to which paper it refers and must be accompanied by a brief abstract. The same publication schedule as for regular articles is followed, and page proofs are sent to authors.

## Comment on "Brownian motion of two interacting particles under a square-well potential"

V. Berdichevsky and M. Gitterman Department of Physics, Bar-Ilan University, Ramat Gan 52900, Israel (Received 1 May 1995)

We present the exact solution of the Fokker-Planck equation for a particle escaping from the square-well potential under the influence of white noise. The mean-square displacement obeys the Einstein relation, in contrast to the conclusion of the recent article by Akio Morita [Phys. Rev. E 49, 3697 (1994)].

PACS number(s): 05.40.+j

Diffusion over potential barriers is an old problem of great importance in physics, the simplest model for such a problem being the escape of a particle from a square well of depth  $V_0$  and width u (Fig. 1) under the influence of white noise. The Fokker-Planck equation for the probability distribution function  $\rho(x,t)$  for the position x of a particle at time t reads [2]

$$\partial_t \rho = \partial_x [\partial_x \rho + A\delta(x - u)\rho]. \tag{1}$$

Here  $A = V_0/T$ , where the temperature T is measured in units of energy and the time is measured in units of length squared (the diffusion coefficient is chosen equal to unity). The initial condition is assumed to be a  $\delta$  function,

$$\rho(x, t = 0) = \delta(x - x_0). \tag{2}$$

The wall at x = 0 was taken as a reflective boundary, i.e., the boundary condition at x = 0, along with finiteness of

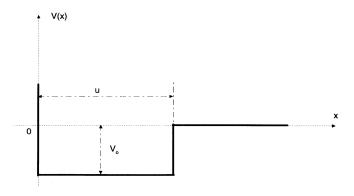


FIG. 1. Square-well potential V(x).

 $\rho$  at  $x \to \infty$ , is

$$\partial_x \rho = 0$$
 at  $x = 0$ ,  $\rho$  is finite at  $x \to \infty$ .

A quite sophisticated analysis of (1)–(3) was performed in Ref. [1] which includes, in particular, the different asymptotic forms of the confluent hypergeometric functions, leading to a very strange result for the asymptotic expansion of the mean-square displacement  $\langle x^2 \rangle$ . Indeed, it was found [1] that the Einstein relation does not hold ( $x^2 \not\approx 2t$  at  $t \to \infty$ ) and the diffusion is anomalous ( $x^2 \sim t^n$  with n > 1). A similar result was obtained in Ref. [1] for the escape from a three-dimensional spherically symmetric potential well. However, we believe the analysis in Ref. [1] is flawed because of the incorrect use of the Laplace transform of Eq. (1); namely, the author of Ref. [1] ignored the existense of the branch point in inverting the Laplace transform. We present here the exact solution of (1)-(3), first for the simple case  $x_0 = u$ , i.e.,  $\rho(x,t=0) = \delta(x-u)$  (initially a particle is located at the right end of the barrier), and then for the general case (2). It is self-evident that the asymptotic results will not depend on the precise initial position of the particle inside a well. In contrast to Ref. [1], we solve Eq. (1) in two regions, 0 < x < u and  $u < x < \infty$ , and then match the solutions obtained at x = u. In each of these regions, Eq. (1) does not contain a  $\delta$  function. Let us perform the Laplace transform,  $\hat{\rho}(s) = \int_0^\infty \rho(t)e^{-st}dt$ :

$$\partial_x^2 \hat{\rho} = s\hat{\rho} - \rho(x,0) = s\hat{\rho} - \delta(x-u), \tag{4}$$

where the initial condition  $\rho(x,t=0)=\delta(x-u)$  has been used. The solutions of Eq. (4) in the two regions of interest are

$$\hat{\rho}_1 = C_1 \cosh(\sqrt{s}x) , \quad 0 < x < u,$$

$$\hat{\rho}_2 = C_2 \exp^{-\sqrt{s}x} , \quad u < x < \infty.$$
(5)

The boundary conditions (3) were used to obtain (5). Turning now to the matching conditions at x=u, one has to take into account the jump of  $\rho(x,t)$  at this point, namely [2],  $\rho(u+\epsilon,t)e^A=\rho(u-\epsilon,t)$ , and the continuity of the current. The latter is assured by integrating (4) over the x region  $(u-\epsilon,u+\epsilon)$  with an infinitesimal  $\epsilon$ . This procedure gives

$$\partial_x \hat{\rho}_2|_{x=u+\epsilon} - \partial_x \hat{\rho}_1|_{x=u-\epsilon} = -1. \tag{6}$$

The two matching conditions for  $\rho$  and  $\partial_x \rho$  at x=u lead to the following values for the constants  $C_1$  and  $C_2$ :

$$C_{1} = \frac{1}{\sqrt{s}[\sinh(\sqrt{s}u) + e^{-A}\cosh(\sqrt{s}u)]};$$

$$C_{2} = \frac{e^{\sqrt{s}u - A}\cosh(\sqrt{s}u)}{\sqrt{s}[\sinh(\sqrt{s}u) + e^{-A}\cosh(\sqrt{s}u)]}.$$
(7)

Equations (5) and (7) yield a complete solution of the time-dependent problem (1) in the Laplace-transformed form. For a comparision with Ref. [1], we restrict ourselves to the calculations performed in Ref. [1]. The Laplace transform of  $\langle x^2 \rangle$  becomes

$$\mathcal{L}\{\langle x^2 \rangle\} = \frac{2}{s^2} + \frac{u^2}{s} + \frac{2u\sqrt{s}\cosh(\sqrt{s}u)(1 - e^A)}{s^2[e^A\sinh(\sqrt{s}u) + \cosh(\sqrt{s}u)]}.$$
 (8)

Expanding the hyperbolic functions in (8) in terms of negative exponentials and then in a series by the binomial theorem, one can perform the Laplace transform exactly [3]:

$$\langle x^{2} \rangle = 2t + u^{2} - \frac{4u}{\sqrt{\pi}} \sum_{n=0}^{\infty} \left( \frac{e^{A} - 1}{e^{A} + 1} \right)^{n+1} \left\{ \sqrt{t} \left[ \exp\left( \frac{-u^{2}(n+1)^{2}}{t} \right) + \exp\left( \frac{-u^{2}(n)^{2}}{t} \right) \right] - \sqrt{\pi}u(n+1)\operatorname{erfc}\left( \frac{u(n+1)}{\sqrt{t}} \right) - \sqrt{\pi}un\operatorname{erfc}\left( \frac{un}{\sqrt{t}} \right) \right\}.$$

$$(9)$$

For A=0, (9) reduces to the well-known field-free result  $\langle x^2 \rangle = 2t + u^2$ . Using the asymptotic expansion of  $\operatorname{erfc}(z)$  for small z,  $\operatorname{erfc}(z) \cong 1 - 2z/\sqrt{\pi}$ , leads to the following asymptotic expansion for  $t \to \infty$ :

$$\langle x^2 \rangle |_{t \to \infty} = u^2 - 2e^A u^2 + 2(e^A)^2 u^2 + 2t$$
  
$$-\frac{4u(e^A - 1)}{\sqrt{\pi}} \sqrt{t} + O\left(\frac{1}{\sqrt{t}}\right). \tag{10}$$

If one considers the general initial condition (2), the result (10) remains unchanged except that one has to replace  $u^2$  in the first term of the right-hand side of (10) by  $x_0^2$ . In addition to  $\langle x^2 \rangle$ , we also calculated the time-dependent probability P(t) to remain inside the well. The solution (5) and (7) yields for the Laplace transform of P(t)

$$\mathcal{L}\lbrace P(t)\rbrace = \int_0^u \hat{\rho}_1(x, s) dx$$

$$= \frac{\sinh(\sqrt{s}u)}{s[\sinh(\sqrt{s}u) + e^{-A}\cosh(\sqrt{s}u)]}.$$
(11)

The inverse Laplace transform of (11) is then

$$P(t) = \frac{e^A}{e^A + 1} \sum_{n=0}^{\infty} \left( \frac{e^A - 1}{e^A + 1} \right)^n \left[ \text{erfc} \left( \frac{un}{\sqrt{t}} \right) - \text{erfc} \left( \frac{u(n+1)}{\sqrt{t}} \right) \right],$$
 (12)

which has the following asymptotic form for  $t \to \infty$ :

$$P(t)|_{t\to\infty} \approx \frac{ue^A}{\sqrt{\pi t}}.$$
 (13)

The same limiting form is obtained for the initial condition (2). Finally, in contrast to Ref. [1], we have found that the Einstein relation does hold for a particle escaping from the potential well shown in Fig. 1, as one would expect for obvious physical reasons.

<sup>[1]</sup> A. Morita, Phys. Rev. E 49, 3697 (1994).

<sup>[2]</sup> H. Risken, The Fokker-Planck Equation (Springer-Verlag, Berlin, 1984).

<sup>[3]</sup> G. Roberts and H. Kaufman, *Table of Laplace Transforms* (Saunders, Philadelphia, 1966).